

Local standard of rest (LSR)

- Local standard of rest is the point that is located today at the present position of the sun and that moves along a perfectly circular orbit around the galactic center. So the locus of LSR is not the actual orbit of the Sun. But this is a good enough reference because although LSR changes with time, the timescale of such changes is very large so periodic corrections can be done when considerable change occurs.
- The need for LSR arises from the fact that galaxy does not rotate as a single rigid entity but has a differential rotation. This means as one goes further away from the apex, the velocity of objects keep decreasing as shown in the figure below
- The velocity of an object with respect to the LSR is called its peculiar velocity and has three components represented as (u,v,w) where u is the velocity towards(positive) or away (negative) from the the G.C , v is the velocity of circular motion within the plane and w is the velocity towards(positive) or away (negative) from the NGP

$$U = \frac{dR}{dt} \quad (1)$$

$$V = \frac{Rd\phi}{dt} \quad (2)$$

$$W = \frac{dz}{dt} \quad (3)$$

- Since the motion of LSR is restricted to the plane , its u and w components are zero . Therefore $U_{LSR} = W_{LSR} = 0, V_{LSR} = V(R_0)$
- Peculiar velocity of any celestial object is $(u,v,w) = (u - U_{LSR}, v - V_{LSR}, w - W_{LSR}) = (u, v - V(R_0), w)$
- If $\vec{\Delta v}$ is the velocity of a star with respect to the sun, the peculiar velocity of the star is

$$\vec{V} = \vec{V}_{sun} + \vec{\Delta v} \quad (4)$$

$$\vec{\Delta v} = \vec{V} - \vec{V}_{sun} \quad (5)$$

$$\langle \vec{\Delta v} \rangle = \left\langle \frac{\vec{v}_1 + \vec{v}_2 + \dots + \vec{v}_n}{n} \right\rangle - \vec{V}_{sun} \quad (6)$$

- Since the average of the random velocities of objects in the solar neighbourhood is zero, the mean of the velocity of stars is the peculiar velocity of the sun

$$\vec{V}_{sun} = -\langle \vec{\Delta v} \rangle \quad (7)$$

- Thus the average velocities of stars within the solar neighbourhood will give us the particular velocity of sun.
- Once both the direction and magnitude of this $\overrightarrow{\Delta v}$ is calculated, each component of the Sun's velocity can be obtained as

$$V_{sun} = (u_{sun}, v_{sun}, w_{sun}) \quad (8)$$

- Where,

$$u_{sun} = -10Km/sec \quad (9)$$

$$v_{sun} = 5Km/sec \quad (10)$$

$$w_{sun} = 7Km/sec \quad (11)$$